

$g_1(x)$ and $g_2(x)$ in the Meson Cloud Model

A. I. Signal

*Institute of Fundamental Sciences PN461
Massey University
Private Bag 11 222, Palmerston North
New Zealand*

Abstract. We calculate the spin dependent structure functions $g_1(x)$ and $g_2(x)$ of the proton and neutron. Our calculation uses the meson cloud model of nucleon structure and includes the effects of kinematic terms which mix transverse and longitudinal spin components. We find small corrections to the nucleon structure functions, however these are significant for the neutron.

The spin dependent structure functions of the nucleon are the subject of much theoretical and experimental interest. As deep inelastic (and other) experiments become more precise it is hoped that it may be possible to make an unequivocal measurement of a higher twist component in the structure function g_2 of the nucleon. This would give new information on the gluon field inside the nucleon, and its relationship with the quark fields.

In order to make such an unequivocal identification it is necessary to understand the relationship between the structure functions g_1 and g_2 . In particular there are leading twist contributions to g_2 which arise from scattering from the components of the meson cloud of the physical nucleon. These contributions are of the order of 10% of the structure function, and need to be taken into account when calculating the twist-2 part of g_2 .

The Meson Cloud Model (MCM) arises from the crucial observation [1] that the contribution of scattering from the pion cloud of the nucleon scales in the Bjorken limit. This implies that the parton distributions of the nucleon are modified via a convolution between the parton distribution of the meson and the momentum distribution of the meson in the proton, viz.

$$\delta q^p(x) = \int_x^1 \frac{dy}{y} f_{p\pi}(y) q^\pi\left(\frac{x}{y}\right). \quad (1)$$

As well as pions, the MCM takes into account scattering from the other baryon plus meson components in the Fock expansion of the wavefunction i.e.

$$|N\rangle_{\text{physical}} = \sqrt{Z}|N\rangle_{\text{bare}} + \sum_{MB} \int dy d^2\mathbf{k}_\perp \phi(y, k_\perp^2) |M(y, \mathbf{k}_\perp); B(1-y, -\mathbf{k}_\perp)\rangle. \quad (2)$$

The other ingredients of the model are the interaction Lagrangians \mathcal{L}_{int} describing the $N \rightarrow BM$ vertices and the form factors for these vertices. The small probability of finding

high mass states in this model leads to quick convergence of the sum over baryon-meson states for structure function calculations.

The MCM has been applied successfully in spin independent DIS, giving a good description of the HERA data on semi-inclusive DIS with a leading neutron [2, 3], and also dijet events with a leading neutron [4, 5]. In addition the MCM gives a good description of the observed violation of the Gottfried sum rule [6, 7].

To extend the model to spin dependent DIS requires the contributions of both pseudoscalar and pseudovector mesons, particularly the ρ meson. The pseudoscalar contributions mainly ‘dilute’ the bare spin dependent pdfs, however the importance of $L \neq 0$ amplitudes in the cloud cannot be ignored [8, 9]. The pseudovector mesons can contribute directly to the spin dependent pdfs. In earlier work we calculated the spin dependent sea distributions $\Delta\bar{u}(x)$, $\Delta\bar{d}(x)$, $s(x)$ and $\Delta\bar{s}(x)$ [10]. Our results are in good agreement with the HERMES data [11].

The structure functions $g_1(x)$ and $g_2(x)$ are dominated by valence rather than sea distributions, so the most important contributions in the MCM are those affecting the valence quarks, which are $N \rightarrow N\pi$ and $N \rightarrow \Delta\pi$, with $\mathcal{L}_{int} = ig_{NN\pi}\bar{\psi}\gamma_5\pi\psi$, $f_{N\Delta\pi}\bar{\psi}\partial_\mu\pi\chi^\mu + \text{h.c.}$ respectively.

At finite Q^2 the spin of the struck hadron from the cloud, in this case either a nucleon or Δ , is not parallel with the initial spin of the target nucleon. This implies, as shown by Kumano and Miyama [12], that both longitudinal and transverse spin structure functions of the cloud hadrons contribute to the observed structure functions. For a spin 1/2 baryon component of the cloud we have

$$\delta g_1(x, Q^2) = \frac{1}{1+\gamma^2} \int_x^1 \frac{dy}{y} \sum_{i=1,2} (-1)^{i+1} [\Delta f_{iL}(y) + \Delta f_{iT}(y)] g_i^B\left(\frac{x}{y}, Q^2\right) \quad (3)$$

$$\delta g_2(x, Q^2) = \frac{1}{1+\gamma^2} \int_x^1 \frac{dy}{y} \sum_{i=1,2} (-1)^i \left[\Delta f_{iL}(y) + \frac{\Delta f_{iT}(y)}{\gamma^2} \right] g_i^B\left(\frac{x}{y}, Q^2\right). \quad (4)$$

where $\gamma^2 = 4x^2m_N^2/Q^2$ and $\Delta f_{iL,T}(y)$ are the differences between spin up and spin down fluctuation functions projected longitudinally along or transverse to the baryon 3-momentum. Similar expressions exist for higher spin components of the cloud [12, 13]. The fluctuations are calculated using standard techniques in time-ordered perturbation theory in the infinite momentum frame [8, 12, 13]. We find that for longitudinal fluctuation functions $\Delta f_{iL}(y)$ the nucleon and Δ contributions are of similar size, with the $s = 3/2$ state of the Δ being important. For transverse fluctuations $\Delta f_{iT}(y)$ the nucleon contributions are much larger than those from the Δ .

In order to estimate the size of the MCM contributions to g_1 and g_2 , we need also to calculate the structure functions of the ‘bare’ hadrons. We use the MIT bag model and the methods developed by the Adelaide group [9, 14] to calculate the spin dependent pdfs. We also add ‘by hand’ a phenomenological $\Delta g(x)$ such that the integral of $g_1^p(x)$ agrees with experiment. The resulting $g_1^p(x)$ and $g_1^n(x)$ give a reasonable description of the experimental data. To calculate the bare $g_2(x)$ we simply use the leading twist Wandzura-Wilczek term

$$g_2(x) = -g_1(x) + \int_x^1 \frac{dy}{y} g_1(y), \quad (5)$$

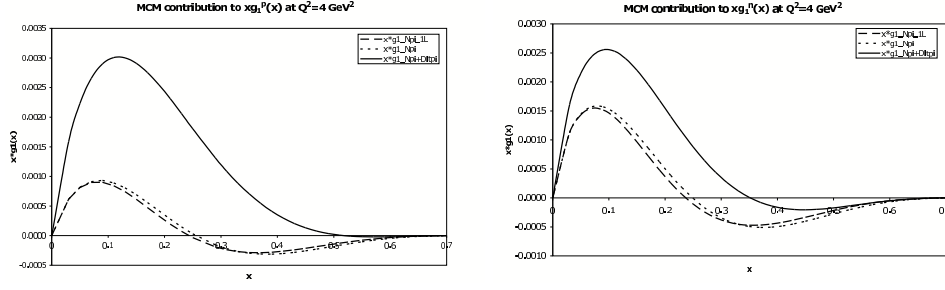


FIGURE 1. Meson cloud Model contributions to g_1 of the proton and neutron. The dashed line is the contribution from longitudinally projected nucleon fluctuations, the dotted line is the total contribution from nucleon fluctuations and the solid line is the total from nucleon and Δ fluctuations.

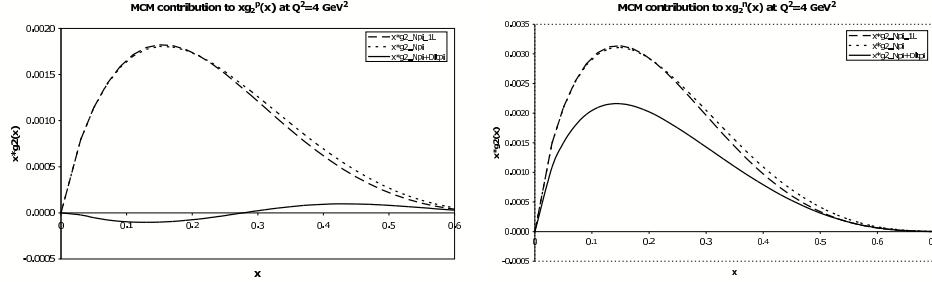


FIGURE 2. Meson cloud Model contributions to g_2 of the proton and neutron. The dashed line is the contribution from longitudinally projected nucleon fluctuations, the dotted line is the total contribution from nucleon fluctuations and the solid line is the total from nucleon and Δ fluctuations.

which also gives a good description of the available experimental data on $g_2^p(x)$ and $g_2^n(x)$.

Our calculations of the contributions to $g_1(x)$ and $g_2(x)$ for both the proton and neutron are shown in figures 1 and 2. For the proton we note that the magnitudes of these contributions are much smaller than the size of the experimental data. We see that the contributions from transversely projected cloud baryons are small, and that the contributions of nucleons and Δ baryons are of similar magnitude, though not necessarily the same sign. For the neutron structure functions the MCM contributions are around 10% of the size of the experimental data. These corrections will be important to consider in any extraction of higher twist components to the neutron structure functions, as they have similar magnitude to these components. Also these corrections have a weak scale dependence which can mimic that of twist-3 contributions at low Q^2 . The correction to $g_2^n(x)$ is positive, and may be able to account for the deviation of the JLab E97-103 data around $x = 0.2$ and $Q^2 = 1 \text{ GeV}^2$ from the Wandzura-Wilczek term [15].

ACKNOWLEDGMENTS

A.S. acknowledges the hospitality and support of the Institute for Particle Physics Phenomenology, Durham University, where portions of this work were done.

REFERENCES

1. J. D. Sullivan, *Phys. Rev. D*, **5**, 1732 (1972).
2. C. Adloff et al., *Eur. Phys. J. C*, **6**, 587 (1999).
3. M. Derrick et al., *Phys. Lett. B*, **384**, 388 (1996); S. Chekanov et al., *Nucl. Phys. B*, **637**, 3 (2002).
4. J. Breitweg et al., *Nucl. Phys. B*, **596**, 3 (2001).
5. A. Aktas et al., hep-ex/0501074, *submitted to Eur. Phys. J. C*.
6. P. Amaudraz et al., *Phys. Rev. Lett.*, **66**, 2712 (1991); M. Arneodo et al., *Phys. Rev. D*, **50**, 1 (1994); M. Arneodo et al., *Phys. Lett. B*, **364**, 107 (1995).
7. E. A. Hawker et al., *Phys. Rev. Lett.*, **80**, 3715 (1998); *Phys. Rev. D*, **64**, 052002 (2001).
8. H. Holtmann, A. Szczurek, and J. Speth, *Nucl. Phys. A*, **569**, 631 (1996).
9. C. Boros and A. W. Thomas, *Phys. Rev. D*, **60**, 074017 (1999).
10. F. G. Cao and A. I. Signal, *Phys. Rev. D*, **68**, 074002 (2003).
11. A. Airapetian et al., *Phys. Rev. Lett.*, **92**, 012005 (2004).
12. S. Kumano and M. Miyama, *Phys. Rev. D*, **65**, 034012 (2002).
13. F. Bissey, F. G. Cao and A. I. Signal, *in preparation*.
14. A. I. Signal and A. W. Thomas, *Phys. Lett. B*, **211**, 481 (1988); *Phys. Rev. D*, **40**, 2832 (1989); A. W. Schreiber, A. W. Thomas, and J. T. Londergan, *Phys. Rev. D*, **42**, 2226 (1990).
15. J. P. Cheng, *presentation at DIS05*.